

Splitting methods for the time dependent Schrödinger equation

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The Schrödinger equation

Consider the autonomous time dependent SE

$$i \frac{\partial}{\partial t} \psi(x, t) = \left(-\frac{1}{2\mu} \nabla^2 + V(x) \right) \psi(x, t)$$

It is separable in its kinetic and potential parts. The solution of the discretised equation is given by

$$i \frac{d}{dt} \mathbf{c}(t) = \mathbf{H} \mathbf{c}(t) \quad \Rightarrow \quad \mathbf{c}(t) = e^{-it\mathbf{H}} \mathbf{c}(0)$$

where $\mathbf{c} = (c_1, \dots, c_N)^T \in \mathbb{C}^N$ and $\mathbf{H} = \mathbf{T} + \mathbf{V} \in \mathbb{R}^{N \times N}$ Hermitian matrix.
Fourier methods are frequently used

$$\begin{aligned} (\mathbf{V}\mathbf{c})_i &= V(x_i) c_i && N \text{ products} \\ \mathbf{T}\mathbf{c} &= \mathcal{F}^{-1} \mathbf{D}_T \mathcal{F} \mathbf{c} && \mathcal{O}(N \log N) \text{ operations} \end{aligned}$$

\mathcal{F} is the fast Fourier transform (FFT)

Consider the Strang-splitting or leap-frog second order method

$$U_2(\tau) \equiv e^{\tau/2V} e^{\tau T} e^{\tau/2V}$$

Notice that $(e^{\tau V} \mathbf{c})_i = e^{\tau V(x_i)} c_i$

the exponentials are computed only once and are stored at the beginning. Similarly, for the kinetic part we have

$$e^{\tau T} \mathbf{c} = \mathcal{F}^{-1} e^{\tau D_T} \mathcal{F} \mathbf{c}$$

This splitting was proposed in:

Feit, Fleck, and Steiger, *J. Comput. Phys.*, 47 (1982), 412.

$$U_n(h) = U_2(\beta_k h) \cdots U_2(\beta_2 h) U_2(\beta_1 h)$$

$$\begin{aligned} U_n(h) &= e^{hb_k \mathbf{B}} e^{ha_k \mathbf{A}} \cdots e^{hb_1 \mathbf{B}} e^{ha_1 \mathbf{A}} \\ &= \Phi_1^*(d_k h) \Phi_1(c_k h) \cdots \Phi_1^*(d_1 h) \Phi_1(c_1 h) \end{aligned}$$

$$U_n(h) = e^{hb_k \mathbf{B}} e^{ha_k \mathbf{A}} \cdots e^{hb_1 \mathbf{B}} e^{ha_1 \mathbf{A}} \quad [\mathbf{B}, [\mathbf{B}, [\mathbf{B}, \mathbf{A}]]] = 0$$

$$U_n(h) = e^{\tilde{\mathbf{B}}(h, \mathbf{b})} e^{ha_k \mathbf{A}} \cdots e^{\tilde{\mathbf{B}}(h, \mathbf{b})} e^{ha_1 \mathbf{A}} \quad [\mathbf{B}, [\mathbf{B}, [\mathbf{B}, \mathbf{A}]]] = 0$$

$$U_n(h) = e^{hb_k \mathbf{H}_0} e^{\epsilon ha_k \mathbf{H}_1} \cdots e^{hb_1 \mathbf{H}_0} e^{\epsilon ha_1 \mathbf{H}_1} \quad \|\epsilon \mathbf{H}_1\| \ll \|\mathbf{H}_0\|$$

$$U_n(h) = e^{hb_k \mathbf{B}} e^{ha_k \mathbf{A}} \cdots e^{hb_1 \mathbf{B}} e^{ha_1 \mathbf{A}} \quad \begin{aligned} &[\mathbf{B}, [\mathbf{B}, [\mathbf{B}, \mathbf{A}]]] = 0 \\ &[\mathbf{A}, [\mathbf{A}, [\mathbf{A}, \mathbf{B}]]] = 0 \end{aligned}$$

$$U_n = U_P U_K U_P^{-1} \implies U_n^p = U_P U_K^p U_P^{-1}$$

Non-autonomous Separable System

Let us now consider the non-autonomous separable system

$$\dot{\mathbf{u}} = (\mathbf{A}(t) + \mathbf{B}(t)) \mathbf{u}$$

with the standard form to convert the system into autonomous

$$\frac{d}{dt} \underbrace{\begin{Bmatrix} \mathbf{u} \\ t_A \\ t_B \end{Bmatrix}}_{\mathbf{y}} = \underbrace{\begin{Bmatrix} \mathbf{A}(t_B)\mathbf{u} \\ 1 \\ 0 \end{Bmatrix}}_{\bar{\mathbf{A}}\mathbf{y}} + \underbrace{\begin{Bmatrix} \mathbf{B}(t_A)\mathbf{u} \\ 0 \\ 1 \end{Bmatrix}}_{\bar{\mathbf{B}}\mathbf{y}}$$

or, equivalently

$$\dot{\mathbf{y}} = (\bar{\mathbf{A}} + \bar{\mathbf{B}}) \mathbf{y}$$

This simple procedure can cause serious drawbacks in the numerical schemes to be used

$$U_n(h) = U_2(\beta_k h) \cdots U_2(\beta_2 h) U_2(\beta_1 h)$$

$$\begin{aligned} U_n(h) &= e^{hb_k \mathbf{B}} e^{ha_k \mathbf{A}} \cdots e^{hb_1 \mathbf{B}} e^{ha_1 \mathbf{A}} \\ &= \Phi_1^*(d_k h) \Phi_1(c_k h) \cdots \Phi_1^*(d_1 h) \Phi_1(c_1 h) \end{aligned}$$

$$U_n(h) = e^{hb_k \mathbf{B}} e^{ha_k \mathbf{A}} \cdots e^{hb_1 \mathbf{B}} e^{ha_1 \mathbf{A}}$$

$$[\mathbf{B}, [\mathbf{B}, [\mathbf{B}, \mathbf{A}]]] = 0$$

$$U_n(h) = e^{\tilde{\mathbf{B}}(h, \mathbf{b})} e^{ha_k \mathbf{A}} \cdots e^{\tilde{\mathbf{B}}(h, \mathbf{b})} e^{ha_1 \mathbf{A}}$$

$$[\mathbf{B}, [\mathbf{B}, [\mathbf{B}, \mathbf{A}]]] = 0$$

$$U_n(h) = e^{hb_k \mathbf{H}_0} e^{\epsilon ha_k \mathbf{H}_1} \cdots e^{hb_1 \mathbf{H}_0} e^{\epsilon ha_1 \mathbf{H}_1}$$

$$\|\epsilon \mathbf{H}_1\| \ll \|\mathbf{H}_0\|$$

$$U_n(h) = e^{hb_k \mathbf{B}} e^{ha_k \mathbf{A}} \cdots e^{hb_1 \mathbf{B}} e^{ha_1 \mathbf{A}}$$

$$\begin{aligned} [\mathbf{B}, [\mathbf{B}, [\mathbf{B}, \mathbf{A}]]] &= 0 \\ [\mathbf{A}, [\mathbf{A}, [\mathbf{A}, \mathbf{B}]]] &= 0 \end{aligned}$$

$$U_n = U_P U_K U_P^{-1} \implies U_n^p = U_P U_K^p U_P^{-1}$$

$$U_n(h) = U_2(\beta_k h) \cdots U_2(\beta_2 h) U_2(\beta_1 h)$$

$$\begin{aligned} U_n(h) &= e^{hb_k B} e^{ha_k A} \cdots e^{hb_1 B} e^{ha_1 A} \\ &= \Phi_1^*(d_k h) \Phi_1(c_k h) \cdots \Phi_1^*(d_1 h) \Phi_1(c_1 h) \end{aligned}$$

$$U_n(h) = e^{hb_k B} e^{ha_k A} \cdots e^{hb_1 B} e^{ha_1 A}$$

~~$$[B, [B, [B, A]]] = 0$$~~

$$U_n(h) = e^{\tilde{B}(h,b)} e^{ha_k A} \cdots e^{\tilde{B}(h,b)} e^{ha_1 A}$$

~~$$[B, [B, [B, A]]] = 0$$~~

$$U_n(h) = e^{hb_k H_0} e^{\epsilon ha_k H_1} \cdots e^{hb_1 H_0} e^{\epsilon ha_1 H_1}$$

~~$$\|\epsilon H_1\| < \|H_0\|$$~~

$$U_n(h) = e^{hb_k B} e^{ha_k A} \cdots e^{hb_1 B} e^{ha_1 A}$$

~~$$\begin{aligned} [B, [B, [B, A]]] &= 0 \\ [A, [A, [A, B]]] &= 0 \end{aligned}$$~~

$$U_n = U_P U_K U_P^{-1} \implies U_n^p = U_P U_K^p U_P^{-1}$$

Problem:
$$H(q, p, t) = \frac{1}{2}p^2 + \frac{1}{2}f(t)q^2 + \varepsilon \sum_{j=1}^s \cos(q - \omega_j t)$$

describes the motion of a charged particle in a magnetic field perturbed by s electrostatic plane waves, each with the same wavenumber and amplitude, but with different temporal frequencies.

Standard split

$$H = \left(\frac{1}{2}p^2 + \frac{1}{2}f(t_B)q^2 + \lambda_A \right) + \left(\varepsilon \sum_{j=1}^s \cos(q - \omega_j t_A) - \lambda_B \right)$$

Our split

$$H = \left(\frac{1}{2}p^2 + \frac{1}{2}f(t_A)q^2 + \lambda_A \right) + \varepsilon \sum_{j=1}^s \cos(q - \omega_j t_A)$$

This split allows to use methods for near-integrable systems in addition to techniques for RKN methods included modified potentials and Processing

SB, F. Diele, C. Marangi, and S. Ragni, **Splitting and composition methods for explicit time dependence in separable dynamical systems.**

JCAM 235 (2010) 646-659.

Polynomial Approximations: Taylor, Chebyshev and Splitting

Let us consider again the linear time dependent SE

$$i \frac{d}{dt} \mathbf{c}(t) = \mathbf{H} \mathbf{c}(t) \quad \Rightarrow \quad \mathbf{c}(t) = e^{-it\mathbf{H}} \mathbf{c}(0)$$

with \mathbf{H} real and symmetric.

This problem can be reformulated using real variables.

Consider $\mathbf{c} = \mathbf{q} + i\mathbf{p}$ then $i \frac{d}{dt} (\mathbf{q} + i\mathbf{p}) = \mathbf{H}(\mathbf{q} + i\mathbf{p})$

Hamiltonian system: $\mathcal{H} = \frac{1}{2} \mathbf{p}^T \mathbf{H} \mathbf{p} + \frac{1}{2} \mathbf{q}^T \mathbf{H} \mathbf{q}$

$$\frac{d}{dt} \begin{Bmatrix} \mathbf{q} \\ \mathbf{p} \end{Bmatrix} = \begin{pmatrix} 0 & \mathbf{H} \\ -\mathbf{H} & 0 \end{pmatrix} \begin{Bmatrix} \mathbf{q} \\ \mathbf{p} \end{Bmatrix}$$

or, in short: $\mathbf{z}' = \mathbf{M} \mathbf{z}$ with $\mathbf{z} = (\mathbf{q}, \mathbf{p})^T$

Formal solution: $\mathbf{O}(t) = \begin{pmatrix} \cos(t\mathbf{H}) & \sin(t\mathbf{H}) \\ -\sin(t\mathbf{H}) & \cos(t\mathbf{H}) \end{pmatrix}$

The Taylor Method

An m -stage Taylor method for solving the linear equation can be written as

$$z_{n+1} = P_m^T(\tau M)z_n$$

where $P_m^T(\tau M)$ is the Taylor expansion of the exponential. It is a polynomial function of degree m which approximates the exact solution up to order m

$$P_m^T(\tau M) \equiv \sum_{j=0}^m \frac{1}{j!}(\tau M)^j = e^{\tau M} + \mathcal{O}(\tau^{m+1})$$

and we can advance each step by using the Horner's algorithm

```
y0 = zn
do i = 1, m
  yi = 1/i * tau * M * yi-1
  zn = zn + yi
enddo
zn+1 = ym
```

This algorithm can be trivially rewritten in terms of the real vectors, **q**, **p**, and it only requires to store two extra vector **q** and **p**.

The Taylor Method

The matrix $P_m^T(\tau M)$ that propagates the numerical solution can be written as

$$P_m^T(\tau H) = \begin{pmatrix} T_1(\tau H) & T_2(\tau H) \\ -T_2(\tau H) & T_1(\tau H) \end{pmatrix}$$

where the entries $T_1(y)$ and $T_2(y)$ are the Taylor series expansion of $\cos(y)$ and $\sin(y)$ up to order m , i.e.

$$P_1^T(x) = \begin{pmatrix} 1 & x \\ -x & 1 \end{pmatrix}, \quad P_2^T(x) = \begin{pmatrix} 1 - \frac{x^2}{2} & x \\ -x & 1 - \frac{x^2}{2} \end{pmatrix}$$

Notice that $\det P_m^T(y) = T_1(y)^2 + T_2(y)^2 \neq 1$
it is not a symplectic transformation! The eigenvalues are given by

$$\lambda_{\pm}^T = T_1 \pm iT_2$$

The scheme is stable if $T_1^2(y) + T_2^2(y) \leq 1$

For practical purposes, we require however

$$T_1^2(y) + T_2^2(y) \leq 1 + tol$$

The Chebyshev Method

The Chebyshev method approximates the action of the exponential on the initial conditions by a near-optimal polynomial given by:

$$u(t) = e^{-itH} u_0 \approx P_m^C(tH) u_0$$

where

$$P_{m-1}^C(tH)u_0 = c_0u_0 + 2 \sum_{k=1}^m c_k T_k \left(\frac{H}{\rho(H)} \right) u_0$$

with $c_k = (-i)^k J_k(t\rho(H))$. Here, $T_k(x)$ is the k th Chebyshev polynomial generated from the recursion

$$T_{k+1}(x) = 2xT_k(x) - T_{k-1}(x), \quad k \geq 1$$

and $T_0(x)=1$, $T_1(x)=x$. $J_k(w)$ are the Bessel functions of the first kind which provides a superlinear convergence for $m > w$ or, in other words, when

$$\frac{t\rho(H)}{m} \leq 1$$

The Chebyshev Method

The Clenshaw algorithm allows to compute the action of the polynomial by storing only two vectors

$$\begin{aligned}d_{m+1} &= 0, & d_m &= 0 \\ \text{do } i &= m - 1, 0 \\ & d_i = c_i z_n + \frac{2}{\rho(H)} H d_{i+1} - d_{i+2} \\ \text{enddo} \\ z_{n+1} &= d_0 - d_2\end{aligned}$$

which can also be easily rewritten in term of the real vectors **q,p**.

The scheme can be written as

$$P_m^C(\tau H) = \begin{pmatrix} C_1(\tau H) & C_2(\tau H) \\ -C_2(\tau H) & C_1(\tau H) \end{pmatrix}$$

As in the Taylor case: $\det P_m^C(y) = C_1(y)^2 + C_2(y)^2 \neq 1$

It is not a symplectic transformation.

The Symplectic Splitting Method

We have built splitting methods for the harmonic oscillator!!!

$$\frac{d}{dt} \begin{Bmatrix} q \\ p \end{Bmatrix} = \left[\underbrace{\begin{pmatrix} 0 & 1 \\ 0 & 0 \end{pmatrix}}_A + \underbrace{\begin{pmatrix} 0 & 0 \\ -1 & 0 \end{pmatrix}}_B \right] \begin{Bmatrix} q \\ p \end{Bmatrix}$$

Exact solution (ortogonal and symplectic)

$$O(t) = \begin{pmatrix} \cos(t) & \sin(t) \\ -\sin(t) & \cos(t) \end{pmatrix}$$

We consider the composition

$$K(h) \equiv \prod_{i=1}^m e^{ha_i A} e^{hb_i B}$$

Notice that

$$e^{hA} e^{hB} = \begin{pmatrix} 1 & h \\ 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 \\ -h & 1 \end{pmatrix} = \begin{pmatrix} 1 - h^2 & h \\ -h & 1 \end{pmatrix}$$

The Symplectic Splitting Method

We have
$$K(h) \equiv \prod_{i=1}^m \begin{pmatrix} 1 - a_i b_i h^2 & a_i h \\ -b_i h & 1 \end{pmatrix} = \begin{pmatrix} K_1 & K_2 \\ K_3 & K_4 \end{pmatrix}$$

$$K_1 = \sum_{i=0}^m k_{1,i} h^{2i} \quad K_2 = \sum_{i=1}^m k_{2,i} h^{2i-1}$$
$$K_3 = \sum_{i=1}^m k_{3,i} h^{2i-1} \quad K_4 = \sum_{i=0}^m k_{4,i} h^{2i}$$

which are polynomials of degree twice higher as in the previous cases for the same number of stages.

The algorithm: a **generalisation** of the Horner's algorithm or the Clenshaw algorithm

```
do i = 1, m
    v = Hp
    q := q + a_i τ v
    v = Hq
    p := p - b_i τ v
enddo
```

It only requires to store one additional real vector of dimension

The Symplectic Splitting Method

The methods preserve symplecticity by construction: $\det K(h) = 1$

Stability: M is stable if $|\text{Tr } K| < 2$, i.e.

$$|K_1 + K_4| = \left| \sum_{i=0}^m (k_{1,i} + k_{4,i}) h^{2i} \right| < 2$$

Theorem: Any composition method is conjugate to an orthogonal method, and unitarity is preserved up to conjugacy.

There is a recursive procedure to get the coefficients of the splitting methods from the coefficients of the matrix K .

We can build different matrices with:

- Large stability domain
- Accurate approximation to the solution in the whole interval (like Chebyshev)
- Methods with different orders of accuracy and very large number of stages.

Taylor **order** **tol=inf** **tol=10-8** **tol=10-4**

10.	0.	0.	0.
15.	0.111249	0.111249	0.111249
20.	0.164515	0.164515	0.164515
25.	0.	0.065246	0.065246
30.	0.	0.108088	0.108088
35.	0.0461259	0.138361	0.322661
40.	0.0804521	0.160884	0.321618
45.	0.	0.178294	0.320804
50.	0.	0.192154	0.32015

Chebyshev **order** **tol=inf** **tol=10-8** **tol=10-4**

20.	0.00362818	0.321723	0.643217
25.	0.00233368	0.384289	0.64029
30.	0.00162599	0.425648	0.638324
35.	0.00119743	0.45502	0.636912
40.	0.000918402	0.476957	0.63585
45.	0.000726646	0.635021	0.635021
50.	0.000589228	0.634356	0.634356
55.	0.	0.633812	0.633812
60.	0.	0.633357	0.633357

Splitting methods

m	r	θ'	y_*/m	$\sum_j (a_j + b_j)$	$\mu_r(\theta'm)$	$\nu_r(\theta'm)$
20	16	1	1.0456	3.0553	0.000611028	0.0258433
30	24	1	1.0246	3.19658	0.0000841871	0.0373544
30	0	1	1.1411	3.04948	$2.91902 \cdot 10^{-13}$	$2.28673 \cdot 10^{-9}$
30	0	0.75	1.027	3.44381	$1.2545 \cdot 10^{-17}$	$5.96706 \cdot 10^{-14}$
40	0	1	1.15953	3.21986	$1.06301 \cdot 10^{-15}$	$1.07587 \cdot 10^{-12}$
30	6	1.4	1.41876	3.0921	0.0000518519	0.0131295

Error bounds

$$\|u_n - u(t)\| \leq \frac{|t| \mu_k(\theta) \|u_0\|_{k+1} + \nu_k(\theta) \|u_0\|_k}{\rho(H)^k}.$$

Schrödinger equation with a Poschl-Teller potential

$$i\frac{\partial}{\partial t}\psi(x, t) = \left(-\frac{1}{2\mu}\frac{\partial^2}{\partial x^2} + V(x) \right) \psi(x, t)$$

with

$$V(x) = -\frac{\alpha^2}{2\mu} \frac{\lambda(\lambda - 1)}{\cosh^2(\alpha x)}$$

$$\mu = 1745, \quad \alpha = 2, \quad \lambda = 24,5,$$

Initial conditions

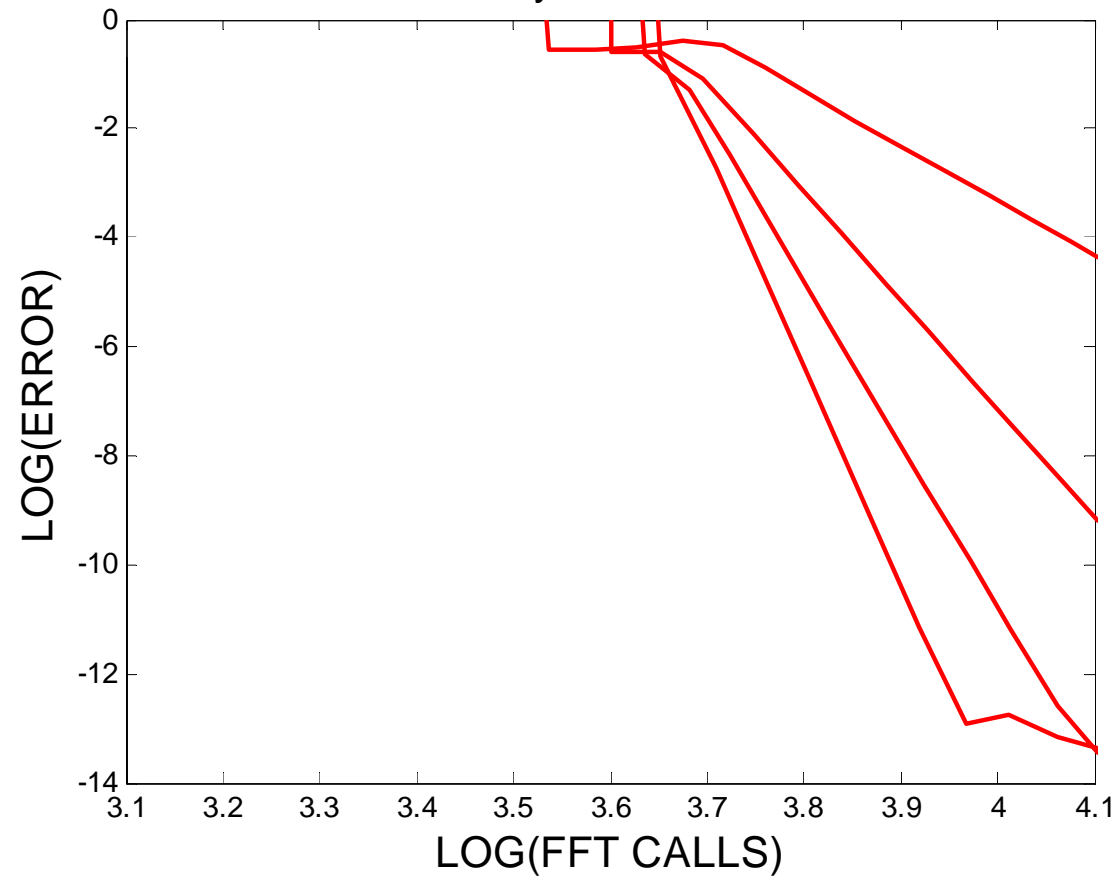
$$\psi(x, 0) = \rho e^{-9^2 x^2}$$

$$t \in [0, 2T] \quad x \in [-5, 5], \quad N = 128 \text{ parts}$$

$$\rho(H) \simeq \frac{1}{2\mu} \left(\frac{\pi}{\Delta x} \right)^2 + (V_{max} - V_{min}) = 1.12$$

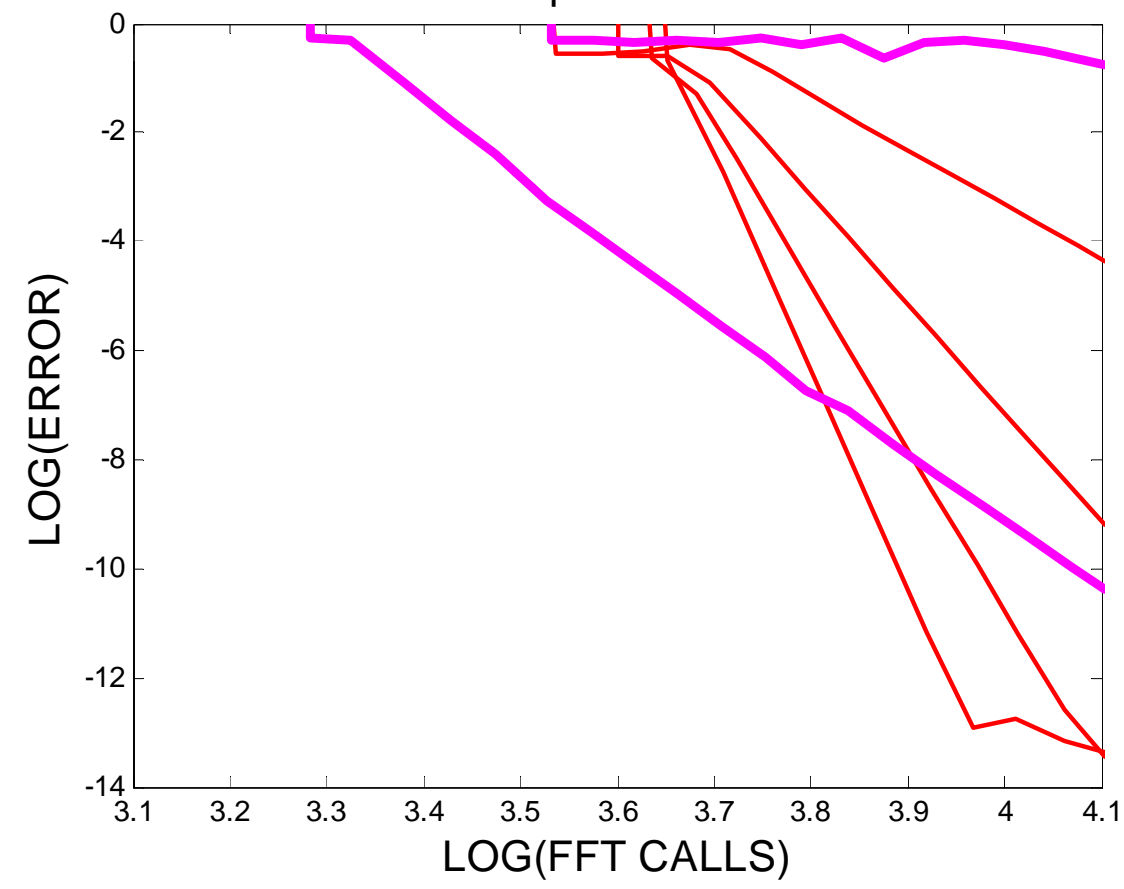
Poschl-Teller potential

Taylor: 10-40



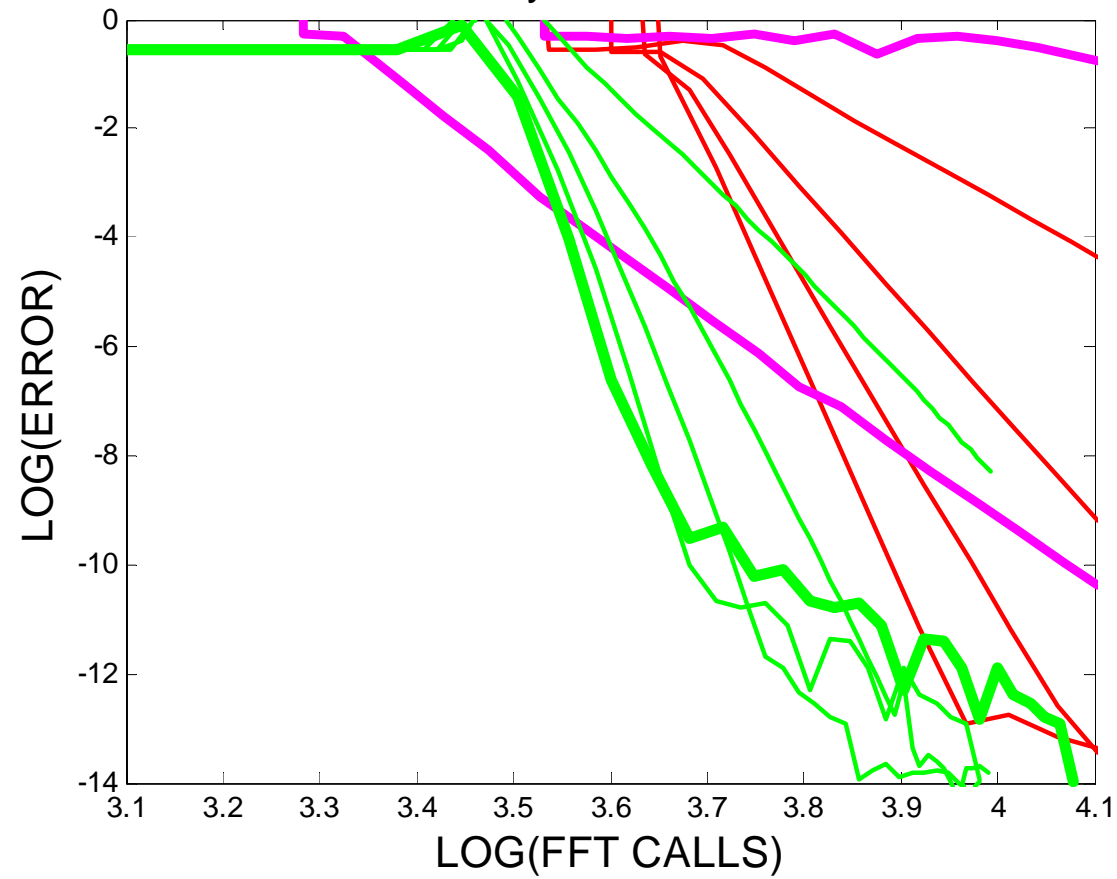
Poschl-Teller potential

Split: 4-12

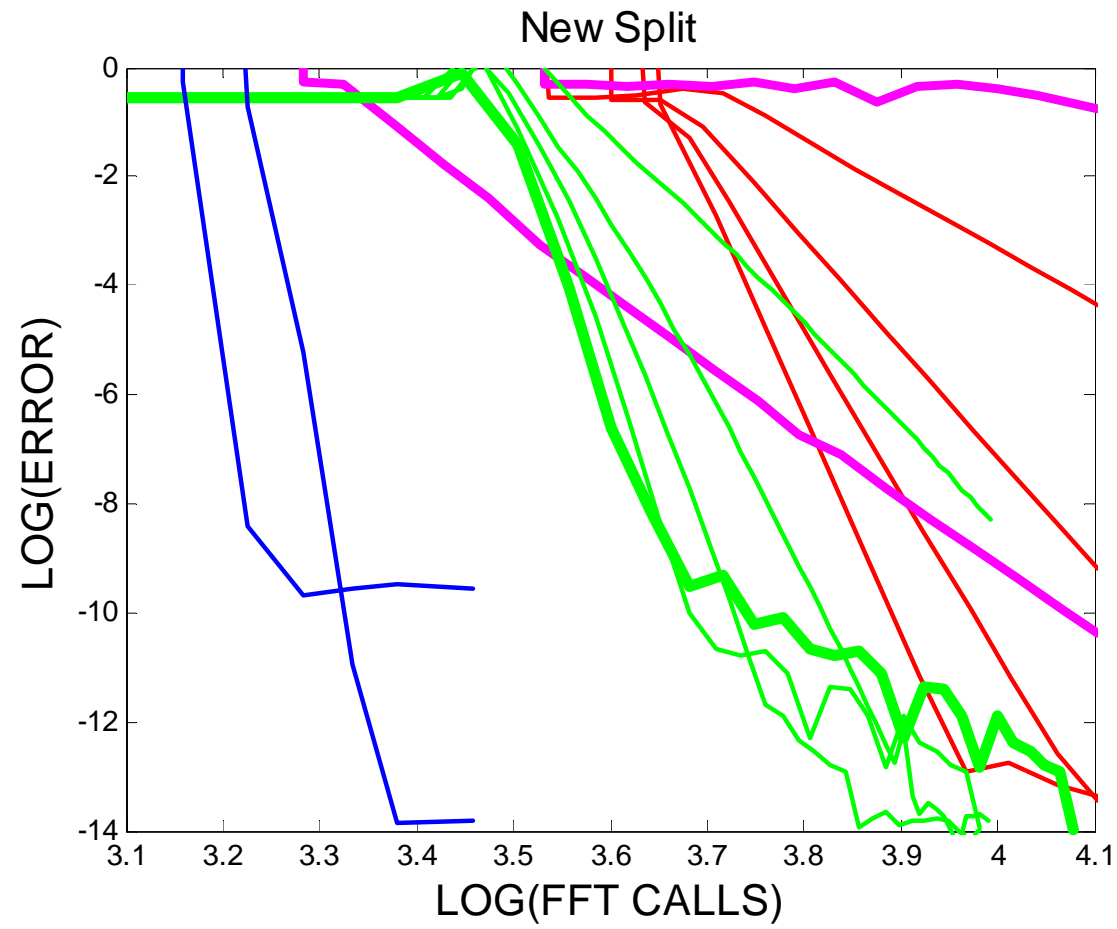


Poschl-Teller potential

Chebyshev: 20-100



Poschl-Teller potential



References

SB, F. Casas, and A. Murua, **Work in Progress.**

Group webpage: <http://www.gicas.uji.es>

SB, F. Casas and A. Murua, **On the linear stability of splitting methods**, Found. Comp. Math., 8 (2008), 357-393.

SB, F. Casas and A. Murua, **Symplectic splitting operator methods for the time-dependent Schrödinger equation**, J. Chem. Phys. 124 (2006) 234105.

Problem: To numerically solve the linear time dependent system

$$x' = M(t)y, \quad y' = -N(t)x$$

It can appear after discretisation of linear PDEs (or their linear part): linear and non-linear Schrödinger equation, Maxwell equations, etc.

The simplest solution is to convert the system into autonomous

$$\begin{aligned} x' &= M(y_t)y & y' &= -N(x_t)x \\ x'_t &= 1 & y'_t &= 1 \end{aligned}$$

where $x_t, y_t \in \mathbb{R}$ The system is no longer linear and the most efficient methods can not be used

$$\begin{aligned} [\mathbf{B}, [\mathbf{B}, [\mathbf{B}, \mathbf{A}]]] &= 0 \\ [\mathbf{A}, [\mathbf{A}, [\mathbf{A}, \mathbf{B}]]] &= 0 \end{aligned}$$

The Magnus series expansion

Theorem (Magnus 1954). Let $A(t)$ be a known function of t (in general, in an associative ring), and let $Y(t)$ be an unknown function satisfying $Y' = A(t) Y$, with $Y(0) = I$. Then, if certain unspecified conditions of convergence are satisfied, $Y(t)$ can be written in the form

$$Y(t) = \exp \Omega(t),$$

where

$$\frac{d\Omega}{dt} = \sum_{n=0}^{\infty} \frac{B_n}{n!} \operatorname{ad}_{\Omega}^n A,$$

and B_n are the Bernoulli numbers.

The convergence is assured if $\|\Omega\| < \pi$

The Magnus expansion

$$X(t) = e^{\Omega(t)} X_0, \quad \Omega = \sum_{k=1}^{\infty} \Omega_k$$

which can be obtained by Picard iteration or using a recursive formula

$$\Omega_1 = \int_0^t A(t) dt$$

$$\Omega_2 = \frac{1}{2} \int_0^t dt_1 \int_0^{t_1} dt_2 [A_1, A_2]$$

$$\Omega_3 = \frac{1}{6} \int_0^t dt_1 \int_0^{t_1} dt_2 \int_0^{t_2} dt_3 \{ [A_1, A_2, A_3] + [A_3, A_2, A_1] \}$$

$$\Omega_4 = \frac{1}{12} \int_0^t dt_1 \int_0^{t_1} dt_2 \int_0^{t_2} dt_3 \int_0^{t_3} dt_4 \{ [A_1, A_4, A_3, A_2] \\ + [A_2, A_3, A_4, A_1] + [A_1, A_2, A_3, A_4] - [A_4, A_3, A_2, A_1] \}$$

where $A_i \equiv A(t_i)$, $[A_1, A_2, A_3] \equiv [A_1, [A_2, A_3]]$

Convergence condition: $\int_0^t \|A(s)\|_2 ds < \pi$

This is a sufficient but
NOT necessary condition

F. Casas, [Sufficient conditions for the convergence of the Magnus expansion](#), J. Phys. A, 40 (2007), 15001-15017

Splitting-Magnus methods

$$z' = (A(t) + B(t))z$$

with

$$A(t) = \begin{pmatrix} 0 & H(t) \\ 0 & 0 \end{pmatrix}, \quad B(t) = \begin{pmatrix} 0 & 0 \\ -H(t) & 0 \end{pmatrix}$$

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We can approximate the solution by the composition

$$\begin{aligned} K(t, h) &= e^{A_m(t, h)} e^{B_m(t, h)} \dots e^{A_1(t, h)} e^{B_1(t, h)} \\ &= \begin{pmatrix} I & M_m \\ 0 & I \end{pmatrix} \begin{pmatrix} I & 0 \\ -N_m & I \end{pmatrix} \dots \begin{pmatrix} I & M_1 \\ 0 & I \end{pmatrix} \begin{pmatrix} I & 0 \\ -N_1 & I \end{pmatrix} \end{aligned}$$

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with the averages

$$M_i = h \sum_{j=1}^k \rho_{ij} H(t + c_j h), \quad N_i = h \sum_{j=1}^k \sigma_{ij} H(t + c_j h)$$

SB, F. Casas and A. Murua, [Splitting methods for non-autonomous Linear systems](#), Int. Jour. Comp. Math., 6 (2007), 713-727

Splitting-Magnus methods

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Splitting-Magnus methods

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with

$$A(t) = \begin{pmatrix} 0 & H(t) \\ 0 & 0 \end{pmatrix}, \quad B(t) = \begin{pmatrix} 0 & 0 \\ -H(t) & 0 \end{pmatrix}$$

This technique can be useful for methods up to order six and for moderate accuracies.

For more accurate results and higher order methods one can use a symmetric second order scheme and the multiproduct methods

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Conclusions

- **Splitting methods are powerful tools for numerically solving the Schrödinger equation**
- **The performance strongly depends on how the system has been split as well as on the choice of the appropriate methods for each problem**
- **Alternatively, one can build methods tailored for particular problems**
- **There are many interesting open problems**